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# Gravity Dual D<sub>3</sub>-braneworld and Closed/Open String Duality

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**ABSTRACT:** A covariantly constant dynamical two-form is exploited on a  $D_3$  -brane to obtain its gravity dual action, governing an  $S^3$  deformed AdS<sub>5</sub> black hole, in a type IIB string theory on  $S^1 \times K3$ . Weinvoke the Kaluza-Klein compactification and identify the T -duality with the open/closed stringduality. Interestingly, the Reissner-Nordstrom black hole is obtained on the "non-Riemannian" braneworld.

### I. INTRODUCTION

In General Theory of Relativity (GTR), the space-time curvatures are based on the intrinsic notion of geometry, by incorporating a nontrivial metric tensor on a (Riemannian) manifold. The fact that the space-time curvatures are governed by the nonlinear matter field, enforces hidden gravity/gaugeduality in Einstein's formulation. The duality becomes evident in presence of a fifth dimension, i.e. in Kaluza-Klein gravity theory. Importantly, the duality takes a precise form in AdS gravity and isestablished as a familiar bulk/boundary or AdS/CFT correspondence [1] in string theory.

In the context, non-perturbative Dirichlet (D-) branes [2] are believed to enlighten the aspects of geometry/matter duality coded in Einstein's formulation. In fact, the nonlinear U (1) gauge dynamics established on a D-brane [3] is indeed remarkable. Along with the higher derivative corrections, the nonlinear gauge sector is believed to govern the space-time curvature. Most importantly, the underlying duality has been considerably explored through the success of 5D Anti de Sitter (AdS<sub>5</sub>)gravity and  $D_3$  -brane gauge theory correspondence established in string theory. Interestingly, thenon-commutative open string (NCOS) leading to a deformed D-braneworld, its gravity correspondence [4] and various (non-linear) charged  $D_3$  -braneworld black holes [5]-[9] have been obtained in the past.

In the paper, we attempt to formalize the nonlinear U (1) gauge dynamics of a Kalb-Ramond (KR)field on a  $D_3$  -brane in terms of its geometric dual in a string theory. We explore a  $D_3$ -braneworld in aplausible fundamental open string formulation and obtain a closed string dual of the  $D_3$ -braneworld, i.e. a curved  $D_3$  -braneworld in 5D. We consider a covariantly constant NS form field to incorporatenew connections, i.e. a torsion, which in turn leads to an unique modified covariant derivative ona vacuum pair of curved braneworld [10]. Our starting point is a flat  $D_3$  -brane, which is knownto govern the gravity decoupling limit  $E \rightarrow E_c = g/(2\pi \alpha')$  in an open string theory.

We revisit the decoupling limit as a cancellation of torsion arising from the zeroth order term in B-field to that of all higher orders, keeping the geometrical notion of Einstein's formulation. In fact, an asymptotic  $D_3$  -brane may be thought of as a transition world-volume connecting the underlying NCOS and afundamental open string theory. As a result, the  $D_3$  -braneworld in the classical regime, i.e.  $g \rightarrow 0$  and  $\alpha' \rightarrow 0$ , may be seen to be described by the nontrivial generic curvatures underlying a(modified) open string theory.



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#### II. KALB-RAMOND FIELD AS A SOURCE OF GRAVITY

#### **NEW CONNECTIONS:**

A dynamical KR field  $B_{\mu\nu}$  incorporates non-trivial connections  $\Gamma_{\mu\nu\lambda}$  on a vacuum pair of gravitationalbrane and antibrane [10, 11, 12, 13, 14, 15]. Since Einstein's gravity decouples, the generic curvature on a curved  $D_3$  -braneworld may be seen to be governed by the "non-Riemannian" geometries. Thevanishing field strength for the dynamical matter introduces the notion of new geometries on the1braneworld. underlying a modified open string theory. Formally, the field strength may be expressed as

$$\mathcal{H}_{\mu\nu\lambda} = (\mathcal{D}_{\mu}B_{\nu\lambda} + F_{\mu\nu}A_{\lambda}) + cyclic (in \,\mu, \nu, \lambda), \qquad (1)$$

where the U (1) field strength  $F_{\mu\nu} = (2\pi\alpha')F_{\mu\nu}$ . The covariant derivative  $\mathfrak{D}_{\mu}$  takes into account non-dynamical metric  $g_{\mu\nu}$  and a dynamical two-form  $B_{\mu\nu}$  and  $\mathfrak{D}_{\lambda}g_{\mu\nu} = 0$ . We rewrite the newconnection as

 $\Gamma_{\mu\nu\lambda} = \Gamma^{(0)}_{\mu\nu\lambda} - \Gamma^{(h)}_{\mu\nu\lambda} + (F_{\mu\nu}A_{\lambda} + cyclic)$ , (2) where  $\Gamma^{(0)}_{\mu\nu\lambda} = -(l/2) (\partial_{\mu}B_{\nu\lambda} + cyclic)$  and  $\Gamma^{(h)}_{\mu\nu\lambda} = -(\Gamma^{\rho}_{\mu\nu}B_{\rho\lambda} + cyclic)$ , respectively, denote the zerothand higher orders, in two-form, defined on a D<sub>3</sub>-braneworld. The gravity decoupling limit, in the formalism, is defined by  $\Gamma^{(0)}_{\mu\nu\lambda} \rightarrow \Gamma^{(h)}_{\mu\nu\lambda}$ . In the low energy regime, the non-Riemannian geometry ona 4D braneworld may be viewed in terms of two equivalent open string descriptions. They are: (i)a matter dominated braneworld for  $|\Gamma^{(0)}_{\mu\nu\lambda}| >> |\Gamma^{(h)}_{\mu\nu\lambda}|$  and (ii) a gravity or geometric braneworld for  $|\Gamma^{(h)}_{\mu\nu\lambda}| >> |\Gamma^{(0)}_{\mu\nu\lambda}|$ . The matter dominated braneworld precisely corresponds to the non-commutativeopen string (NCOS) description. On the other hand, the geometric braneworld may be seen togovern by a (modified) fundamental open string. The underlying open string, in an Einstein gravitydecoupled braneworld, is a new phenomenon in the formalism.

#### **GRAVITY AND GAUGE CURVATURES ON THE BRANE-WORLD:**

The generic gravity and gauge curvatures on the  $D_3$ -braneworld may be computed from the commutator of appropriate gauge covariant derivatives on it. In presence of a gauge field, the appropriate covariant derivative becomes  $D\mu = (\mathfrak{D}\mu - A_{\mu})$ . A priori, the nontrivial curvatures on a  $D_3$ -braneworld may be given by

$$[D_{\mu}, D_{\nu}]A_{\lambda} = \mathcal{F}_{\nu\mu}A_{\lambda} + K_{\mu\nu\lambda}{}^{\rho}A_{\rho} + T^{\rho}{}_{\mu\nu}\mathcal{D}_{\rho}A_{\lambda} + \frac{1}{2}(\Gamma^{\rho}{}_{\mu\sigma}T^{\sigma}{}_{\nu\lambda} + \Gamma^{\sigma}{}_{\nu\lambda}T^{\rho}{}_{\mu\sigma} - \Gamma^{\sigma}{}_{\mu\lambda}T^{\rho}{}_{\nu\sigma} - \Gamma^{\rho}{}_{\nu\sigma}T^{\sigma}{}_{\mu\lambda}A_{\rho}).$$
(3)

The non-linear U (1) field strength  $\mathcal{F}_{\mu\nu} = F_{\mu\nu} + T_{\mu\nu}^{\ \rho}A_{\rho}$ , where  $T_{\mu\nu}^{\ \rho} = -2\Gamma_{\mu\nu}^{\ \sigma}$  is a torsion on thebraneworld. The non-Riemannian curvature tensor is precisely due to the two-form and takes an identical form to the Riemannian-Christoffel curvature tensor  $R_{\mu\nu\lambda\rho}$ . It is given by

$$K^{\rho}_{\mu\nu\lambda} = \partial_{\nu}\Gamma^{\rho}_{\mu\lambda} - \partial_{\mu}\Gamma^{\rho}_{\nu\lambda} + \Gamma^{\sigma}_{\mu\lambda}\Gamma^{\rho}_{\nu\sigma} - \Gamma^{\sigma}_{\nu\lambda}\Gamma^{\rho}_{\mu\sigma}(4)$$

Importantly for a constant torsion,  $K_{\mu\nu\lambda\rho} \rightarrow R_{\mu\nu\lambda\rho}$ . Incorporating the irreducible generic non-Riemannian curvature tensor, we rewrite the D<sub>3</sub>-braneworld action as

$$S = \frac{1}{a} \int d^4x \sqrt{-g} \left( K - \frac{a}{4} F^2 \right).$$
 (5)

A priori, simple looking braneworld may be seen to be governed by the intrinsically coupled gravityand gauge curvatures from the  $B_{\mu\nu}$  and  $A_{\mu}$  equations of motion. They are, respectively, given by

$$\partial_{\lambda} \Gamma^{i\mu\nu} - \frac{1}{2} (g^{\alpha\beta} \partial_{\lambda} g_{\alpha\beta} \Gamma^{i\mu\nu}) = 0$$
  
and  
$$a(\nabla_{\lambda} F^{i\mu}) - 6F_{\lambda\nu} \Gamma^{i\nu\mu} + 12 \nabla_{\lambda} (\Gamma^{i\mu\nu} A_{\nu}) = 0.(6)$$

#### **REISSNER-NORDSTROM ON A PAIR OF BRANE/ANTI-BRANE:**

In this section, we address a non-Riemannian formulation leading to Reissner-Nordstrom black holegeometry on the  $D_3$ braneworld. The anstaz for a two-form and a gauge field leading to a chargedblack hole may be given by  $B_{0\phi} = -M \cos \theta$  and  $A_0 = Q_e/r$ , where M and  $Q_e$  are constants and shallbe identified with the mass and the electric charge. Thegauge choice leads to a nontrivial connection  $\Gamma_{0\phi\phi} = M/2\sin \theta$  and an extremal charged black hole geometry may be seen to emerge through an effective metric  $G_{\mu\nu} = g_{\mu\nu} - B_{\mu\lambda}g^{\lambda\rho}g_{\rho\nu}$  description in an open string theory. However, we consider (generalized) geodesic expansion for the effective metric components  $G_{00} = G^{rr}$ , around an S<sup>2</sup>-symmetric vacua, in a non-relativistic limit. The  $G_{00}$  component of the effective metric may beobtained from a geodesic expansion represented by the derivative corrections of gauge potentials of increasing rank in each term. It is given by  $G_{00} = (g_{00} + h^{(0)}_{00} + h^{(1)}_{00} + h^{(2)}_{00} + \dots)$ . The nontrivial scalarquantities in 2nd, 3rd and 4th terms are essentially the distinct contributions, respectively, out of thescalar, vector and two-form fields. To avoid an ambiguous form of  $h^{(2)}_{00}$ , we rather consider a



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Poincaredual of B-field, i.e. an axion field  $\chi(x)$  as a source of Newtonian potential on the braneworld. Theaxion and gauge gravitational potentials are, respectively, worked out to yield  $h^{(0)}_{00} = -2M/r$  and  $h^{(1)}_{00} = Q_e^2/r^2$ . The axion source on the D3 -braneworld leads to an electric charged black hole which inturn may be viewed in terms of both the charges  $(Q^2 = Q_e^2 + Q_m^2)$  in a boosted coordinate system. In the case, the "non-Riemannian" D<sub>3</sub>-braneworld may be seen to describe a typical Reissner-Nordstromblack hole geometry known in Riemannian geometry. The nontrivial vacua common to two distinct ensor formulations of gravity in string theory is remarkable. It may provide clues to enhance ourunderstanding of de Sitter vacua.

#### III. **GRAVITY DUAL D3-BRANEWORLD**

The gravity and gauge curvatures on the curved  $D_3$ -braneworld in 4D may be seen to have itsorigin in a five dimensional gravity theory governed by a dynamical B-field in a truncated closedstring theory. A priori, one may begin with the type IIB string theory on  $S^1 \times K3$ . We consider anon-dynamical metric and for simplicity, focus on constant scalars and gauge fields. The genericthe notion of closed string is maintained in 5D by the (covariantly constant) dynamical B-field in absence of the dynamical metric on the curved braneworld. The relevant closed string action maybe given by

$$S = \int d^5 y \sqrt{-\tilde{g}} \tilde{K} \quad , \tag{7}$$

where  $\tilde{g}_{uv}$  is a constant metric in the Cartesian coordinates. An ansatz for the B-field, around as<sup>3</sup>- symmetric vacua, is worked out to yield

 $B_{\theta\psi} = B_{r\psi} = r_0$ ,  $B_{\theta\psi} = q(\cot\theta\sin^2\psi)$  and  $B_{\psi\phi} = q(\cos\theta\sin^2\psi)$ , (8) where  $0 < \theta \le \pi$ ,  $0 < \psi \le \pi$ ,  $0 \le \phi \le 2\pi$  and (r<sub>0</sub> and q) are are arbitrary constants. The nontrivial connections are given by  $\Gamma_{\theta\psi\phi} = (q/2)(\sin\theta\sin^2\psi)(9)$ 

Importantly, the D-braneworld tension turns out to be a significant vacuum energy density in theeffective metric description. As a result, the radial coordinate picks up an asymptotic cut-off b, theAdS radius, in the curved Dbraneworld formalism. The 5D black hole metric may be constructed using the B-field ansatz around an AdS vacua. The  $AdS_5$  geometry may be seen to be described by an  $AdS_2$ , and an  $S^2$ , deformed  $S^3$ -symmetry in the closed string theory. It is given by

$$ds^{2} = -\left(1 + \frac{r^{2}}{b^{2}} - \frac{r_{0}^{2}}{r^{2}}\right)dt^{2} + \left(1 + \frac{r^{2}}{b^{2}} - \frac{r_{0}^{2}}{r^{2}}\right)^{-1}dr^{2} + r^{2}\left(1 + \frac{2f^{2}}{r^{4}}\right)d\Omega_{3}^{2} - \frac{B_{\theta\psi}^{2}}{r^{2}}d\Omega_{2}^{2} , \quad (10)$$

where  $f = q(\cot \theta \sin \psi)$ . A priori, the fundamental closed string action (7), may alternately beviewed in an effective metric description to yield the usual Kaluza-Klein gravity in presence of acosmological constant. In that case, the geometry may be seen to be governed by the SchwarzschildAdS<sub>5</sub> black hole with  $S^3$ -symmetry [16, 17]. In other words, AdS<sub>5</sub> black hole in the B-field theory of gravity governs an deformed S<sup>3</sup>, where as the 3D-spherical symmetry is preserved in the metric sourcetheory of gravity. As a result, the  $D_3$  -braneworld action (7) possibly governs a generalized gravitydynamics, which is beyond a typical 5D Kaluza-Klein gravity theory. For a constant  $\Gamma_{\theta\psi\phi}\neq = 0$ , the conformal factor and the  $S^3$  deformation geometry become trivial in the AdS<sub>5</sub> black hole (10). In the case, the Reimannian notion of manifold becomes manifest and the geometry is in agreement with the Schwrazschild  $AdS_2 \times S^3$ black hole obtained in the Kaluza-Klein theory. The Kaluza-Klein compactification of the fifth dimension in the closed string theory (7), in a tatic gauge, precisely reduces to the gravity dual  $D_3$  -brane action (5). For instance, if  $\psi$  denotes acompactified coordinate, the gauge potential may be worked out in 4D, from the B-field ansatz, toyield

 $A_{\mu} = (r_0, r_0, q \cot \theta, -q \cos \theta).(11)$ 

The linear electromagnetic field, in the gravity dual  $D_3$ -braneworld, is a priori governed by the adial component of the magnetic field  $B_r = q/r^2$ , along with a nontrivial B-field in a Kaluza-Klein compactified closed string theory<sup>5</sup>. It is thought provoking to understand the Kaluza-Kleincompactification in a B-field theory of gravity as a T -duality between a, fundamental, open/closedstring. It may also be understood as a (bulk) gravity/(boundary) gauge theory correspondence. In fact, the  $D_3$  -braneworld action possesses a dual notion in a type IIB string theory. They are: (i) afundamental closed string description in 5D and (ii) a fundamental open string in 4D. The geometricbraneworld curvature in 4D essentially confirms the fundamental nature of open string there.

Interestingly, all the higher derivative corrections vanish for a self-dual electro-magnetic field [7]. As a result, the deformed AdS 5 (10) in a closed string theory may be seen to reduce to an AdS<sub>4</sub>black hole whose curvature is due to the nonlinear electric/magnetic charge (q) in an open stringtheory obtained by one of the author in ref.[9].



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#### IV. CONCLUDING REMARKS AND RESULTS

The gravity dual of a  $D_3$ -brane action is obtained in a type IIB string theory on  $S^1 \times K3$ . Thegeometrical description on the braneworld confirms a dual nature of two-form, i.e. a closed stringdescription in 5D and a NCOS description in 4D. The fact is reassured by its underlying deformedAdS black holes geometries both in closed and open string theories. As a bonus, we obtain afundamental open string theory T -dual to a closed string theory in 5D. Our analysis leads to thebirth of a generic curvature tensor  $K_{\mu\nu\lambda\rho}$  which retains the properties of Riemann-Christoffeltensorexcept for the symmetric property under the exchange of its pair of indices. However for a constanttorsion, the gravity dual braneworld may be seen to be described by the Riemannian geometry.

A constant two-form ansatz may be obtained for q = 0 in the gravity dual  $D_3$  -braneworld at the expense of an effective dynamical theory of gravity. It restores the  $S^3$  -symmetry and describes an AdS<sub>5</sub> Schwarzschild black hole in a typical Kaluza-Klein gravity in presence of a cosmological constant. On the other hand, the global mode of a B-field in an open string theory is known to yield the deformation geometry [3]. In fact, it may be interesting to investigate the thermal aspects associated with the non-commutative geometry in the curved D-braneworld along the spirit of entropic force of gravity discussed recently [18]. It confirms that the deformation geometry, in string theory, is intrinsic to the B-field source of gravity, which is unlike to a metric source of gravity. In other words, the open/closed string duality in presence of B-field retains the notion of the deformation geometry. Interestingly, the torsion does not break the S<sup>3</sup> -symmetry in the AdS<sub>5</sub> black hole (10), rather it is broken solely by the B-field.

<sup>5</sup>The Kaluza-Klein compactification may also be seen to yield an irreducible curvature precisely due to a generic nonlinear U (1) gauge field in the theory.

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